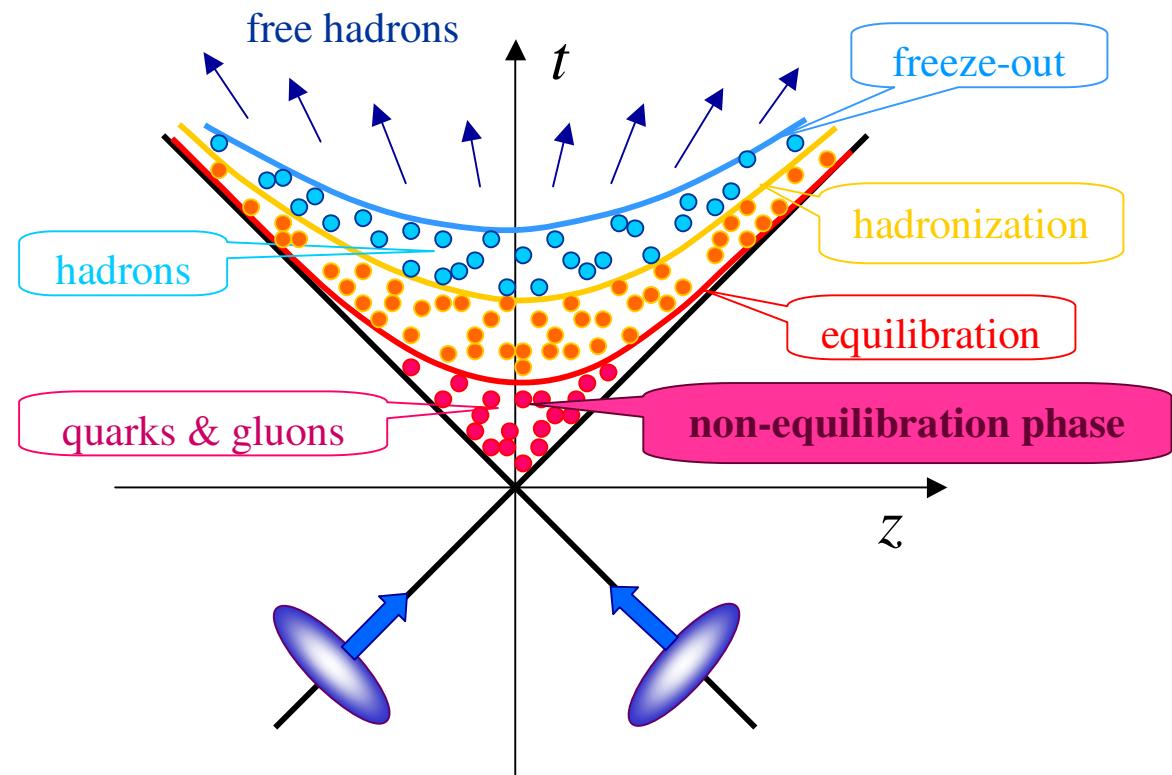
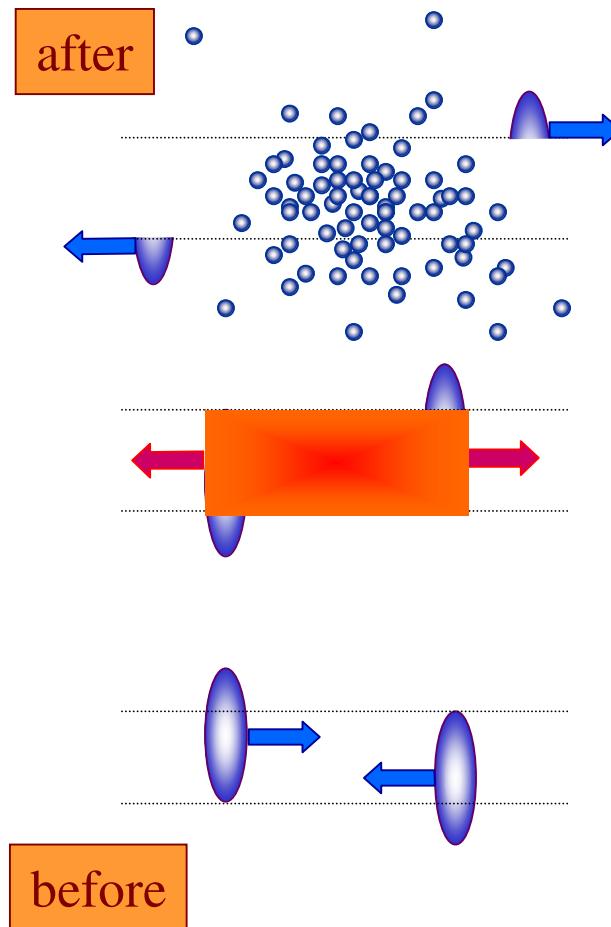


Color Instabilities in Relativistic Heavy-Ion Collisions

Stanisław Mrówczyński

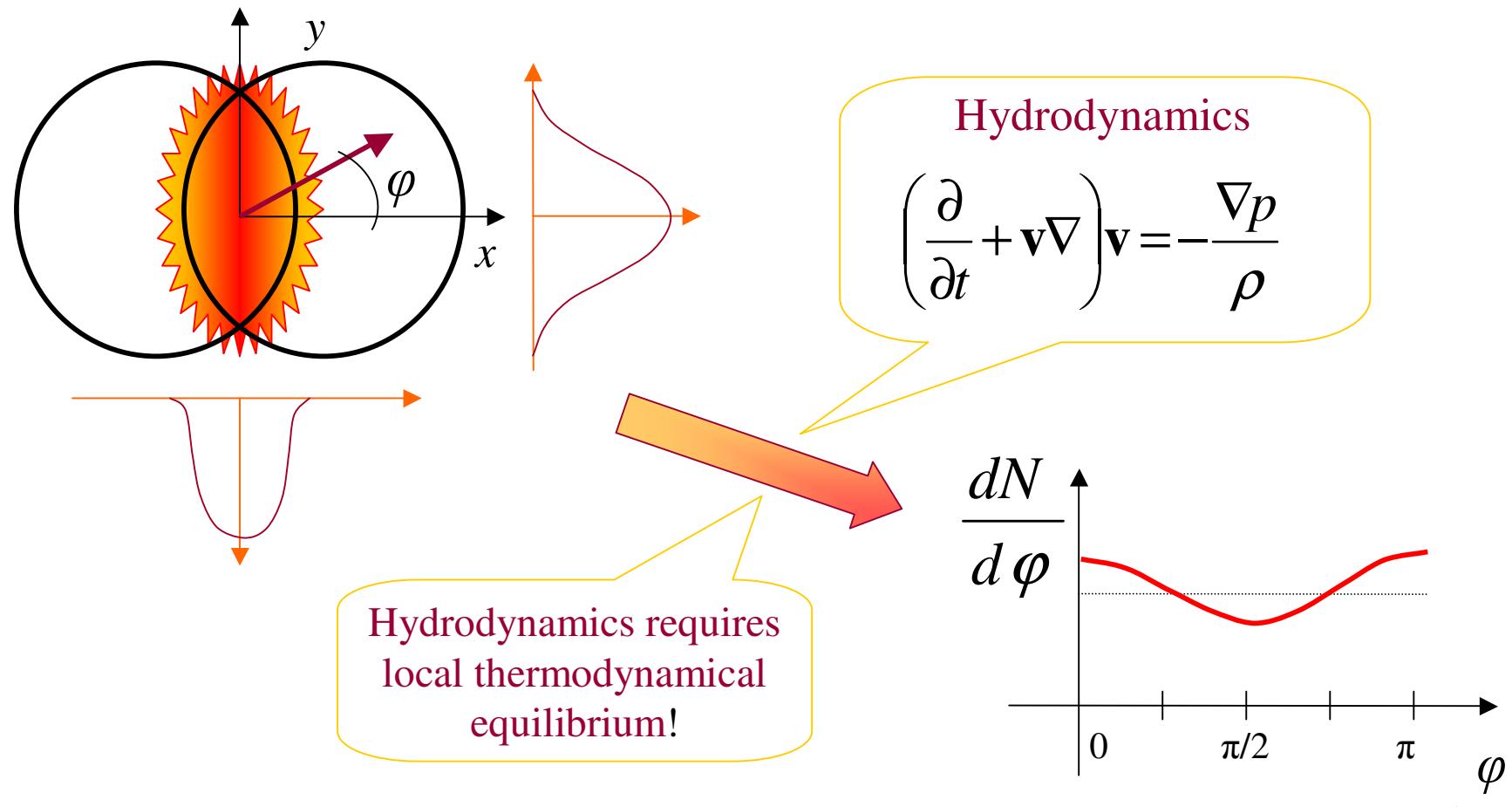
*Jan Kochanowski University, Kielce, Poland
& Institute for Nuclear Studies, Warsaw, Poland*

Course of relativistic heavy-ion collisions



Early stage equilibration @ RHIC

Success of hydrodynamic models in describing elliptic flow



Equilibration is fast

$$v_2 \sim \epsilon = \left\langle \frac{x^2 - y^2}{x^2 + y^2} \right\rangle$$

Eccentricity decays due to the free streaming!

$$\epsilon \searrow \Rightarrow v_2 \searrow$$



$$t_{\text{eq}} \leq 1 \text{ fm}/c$$

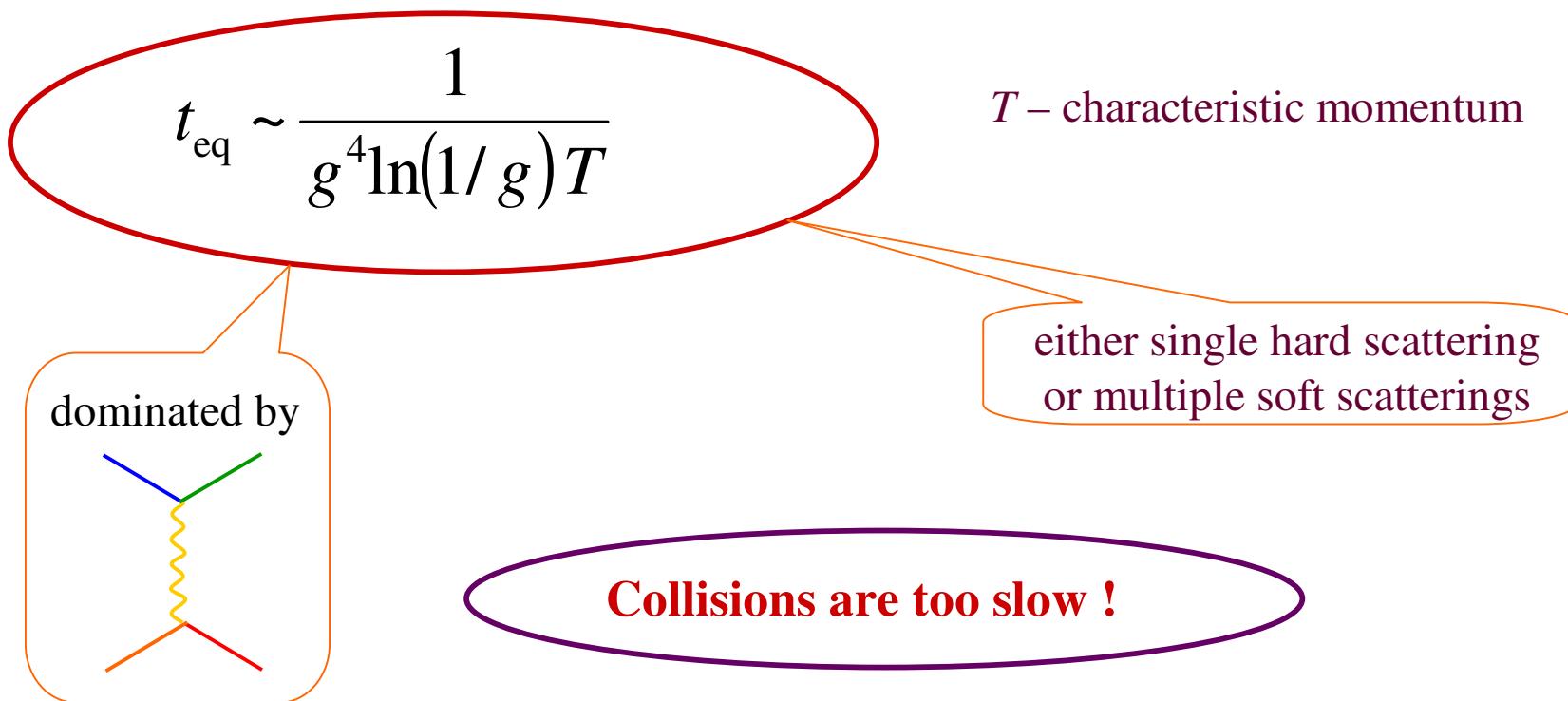
time of equilibration

Collisional equilibration

If QGP is weakly coupled:

$$\alpha_s \equiv \frac{g^2}{4\pi} \ll 1 \text{ -- QCD coupling constant}$$

Time of equilibration driven by parton-parton collisions



Possible conclusion & objection

QGP is strongly coupled: QGP → sQGP

But . . .

asymptotic freedom @ high-energy density

$$\dots \alpha_s \approx 0.3$$

FAQ

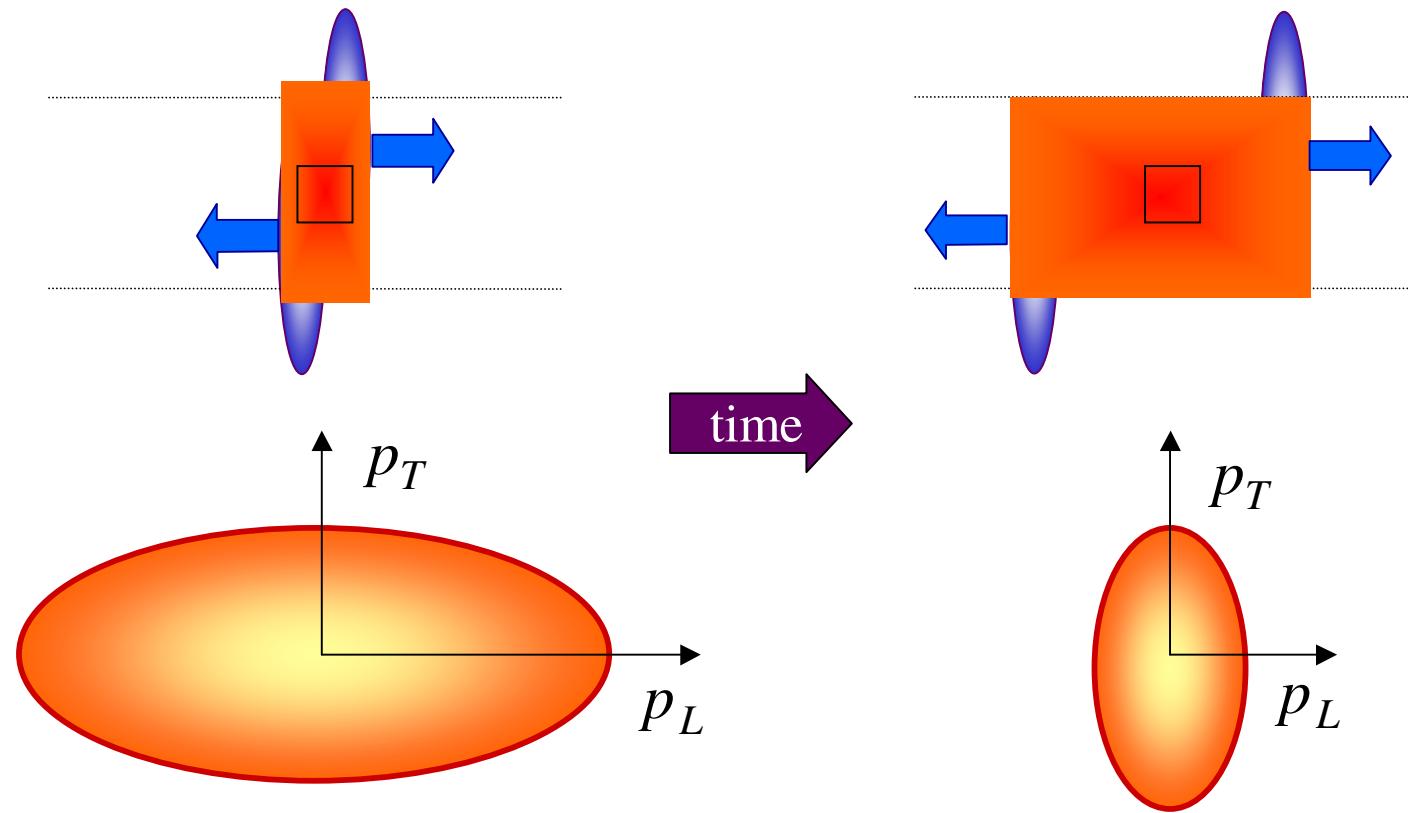
Q: Can weakly coupled QGP equilibrate fast?

A: Yes, due to chromomagnetic instabilities!

Chromomagnetic instabilities

The instabilities occur due to anisotropy of the momentum distribution

Parton momentum distribution is initially anisotropic

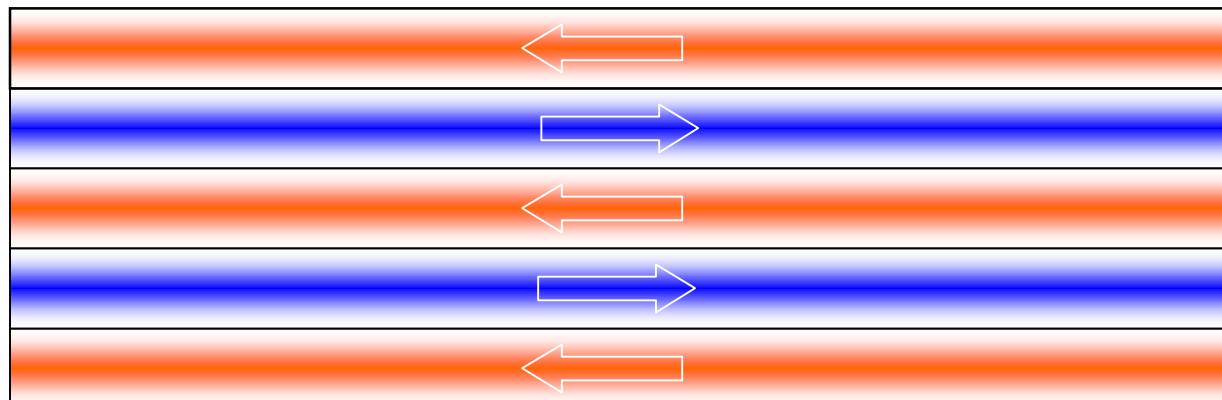


Seeds of instability

$\langle j_a^\mu(x) \rangle = 0$ **but current fluctuations are finite**

$$\langle j_a^\mu(x_1) j_b^\nu(x_2) \rangle = \frac{1}{2} \delta^{ab} \int \frac{d^3 p}{(2\pi)^3} \frac{p^\mu p^\nu}{E_p^2} f(\mathbf{p}) \delta^{(3)}(\mathbf{x} - \mathbf{v}t) \neq 0$$

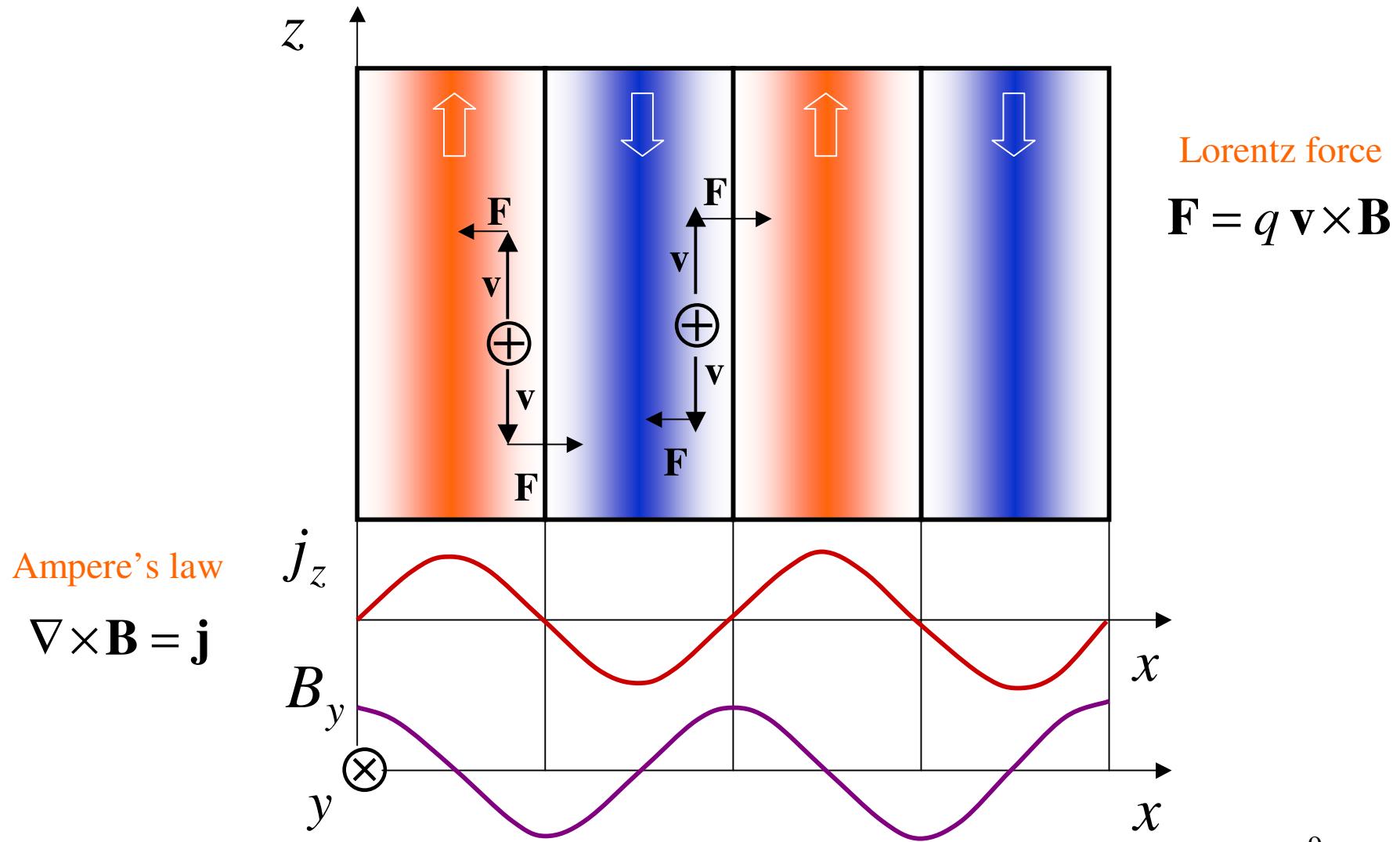
$$x_1 = (t_1, \mathbf{x}_1), \quad x_2 = (t_2, \mathbf{x}_2), \quad x = (t_1 - t_2, \mathbf{x}_1 - \mathbf{x}_2)$$



Direction of the momentum surplus



Mechanism of filamentation

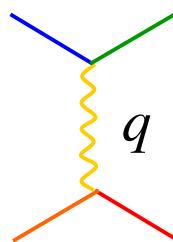


Instabilities are fast

Time scale of processes driven by parton-parton scattering

$$t_{\text{hard}} \sim \frac{1}{g^4 \ln(1/g) T}$$

$$t_{\text{soft}} \sim \frac{1}{g^2 \ln(1/g) T}$$



hard scattering: $q \sim T$

soft scattering: $q \sim gT$

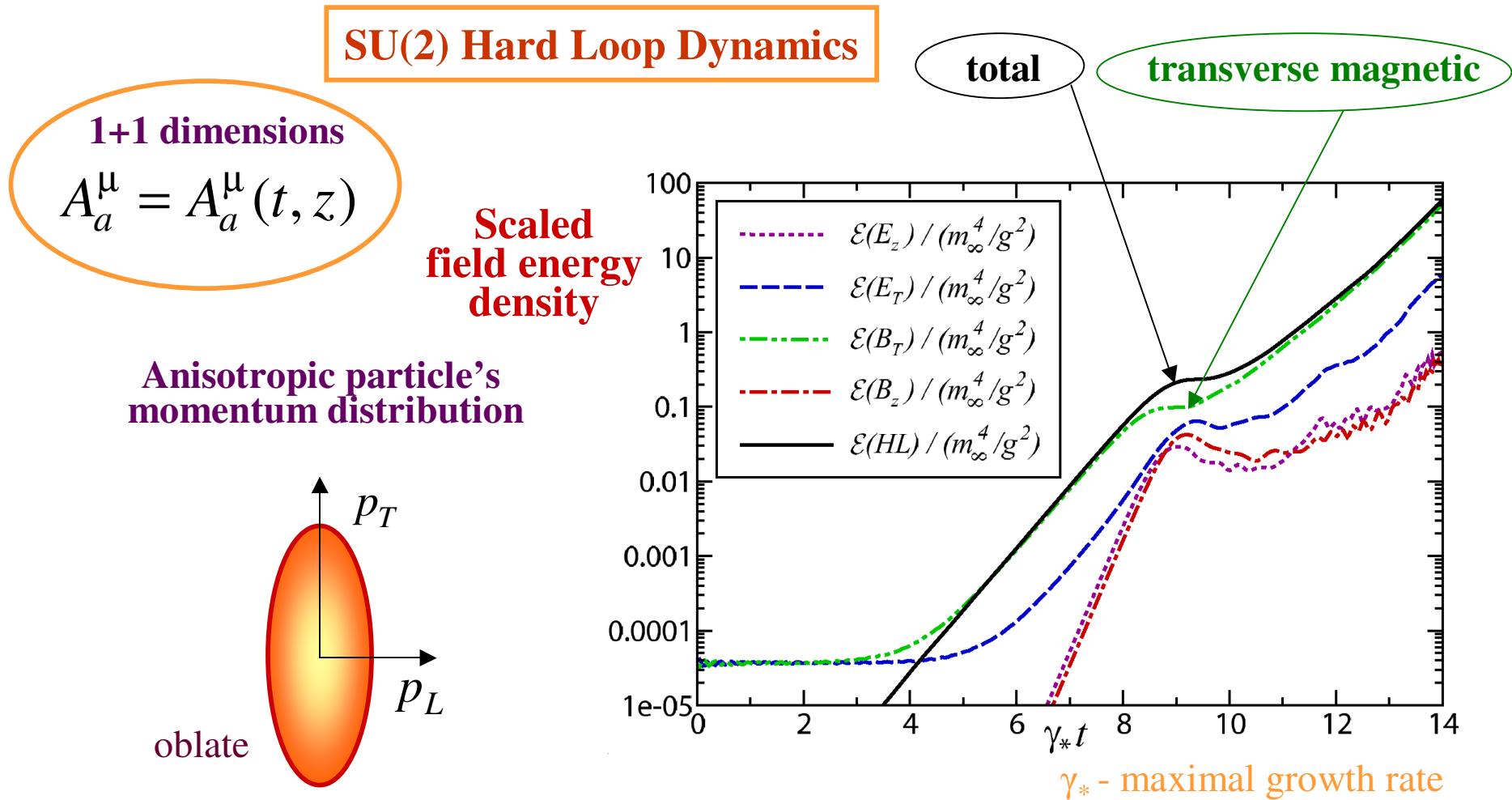
Time scale of collective phenomena

$$t_{\text{collective}} \sim \frac{1}{g T}$$

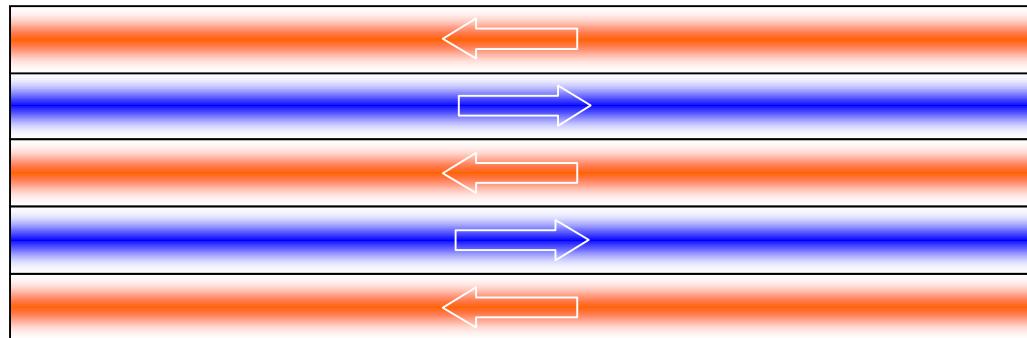
$$g^2 \ll 1 \Rightarrow t_{\text{hard}} \gg t_{\text{soft}} \gg t_{\text{collective}}$$

The instabilities are fast!

Growth of instabilities – 1+1 numerical simulations



Isotropization



Direction of the momentum surplus



$$\begin{array}{c} \mathbf{j}, \mathbf{E} \\ \leftrightarrow \\ \otimes \odot \\ \mathbf{B} \end{array}$$

momentum change
of particles

$$\Delta \mathbf{p} = \int dt \mathbf{F}$$

momentum of fields

$$\mathbf{P}_{\text{fields}} \sim \mathbf{B}^a \times \mathbf{E}^a \sim \mathbf{k}$$

Isotropization – numerical simulation

Classical system of colored particles & fields

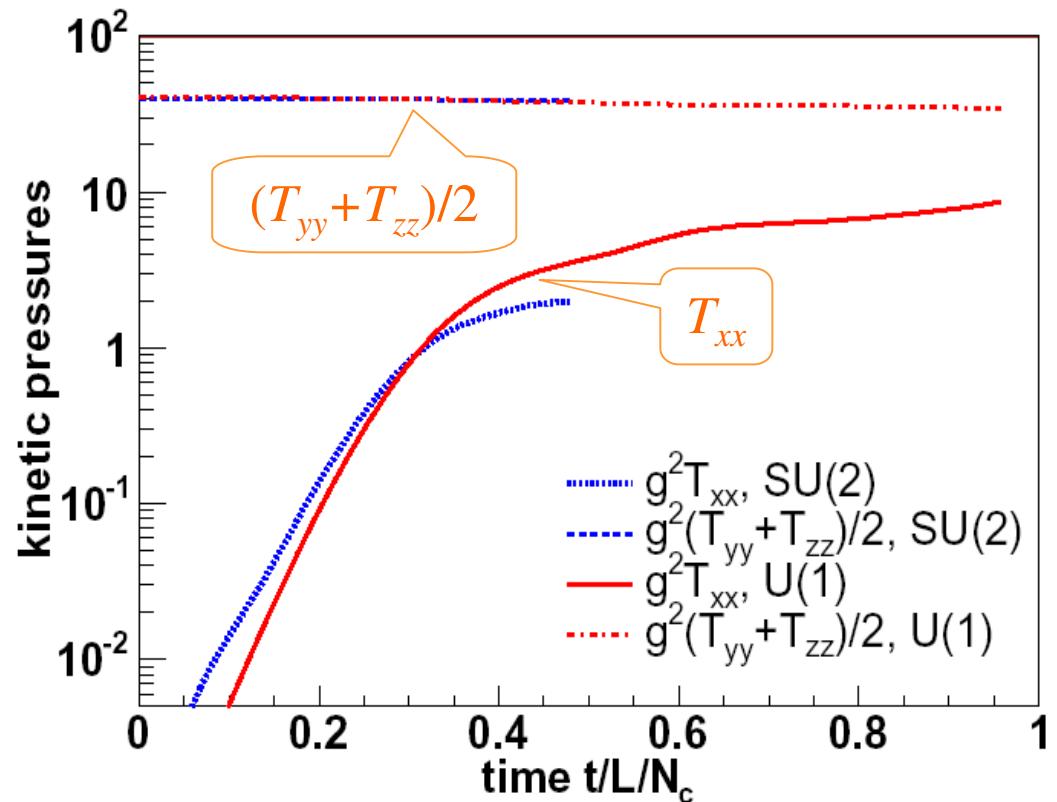
$$T_{ij} = \int \frac{d^3 p}{(2\pi)^3} \frac{p_i p_j}{E} f(\mathbf{p})$$

Initial anisotropy:

$$T_{xx} = 0$$

Isotropy:

$$T_{xx} = (T_{yy} + T_{zz})/2$$



Conclusion

The scenario of instabilities driven equilibration provides a plausible solution to the fast equilibration problem of weakly coupled plasma